

Computer Graphics 13 - Global illumination 2

Tom Thorne

Slides courtesy of Taku Komura
www.inf.ed.ac.uk/teaching/courses/cg

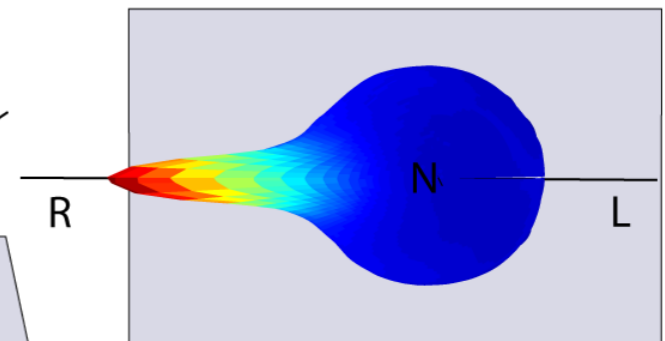
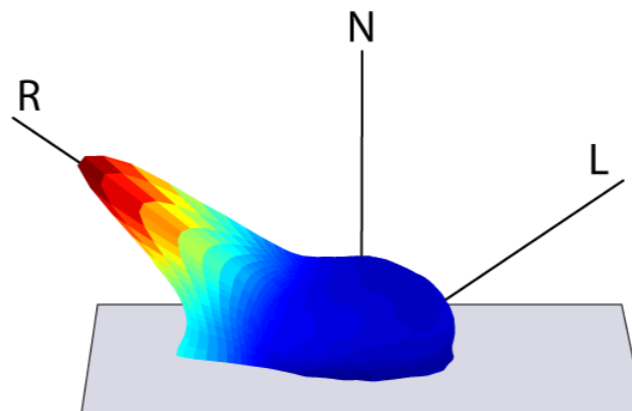
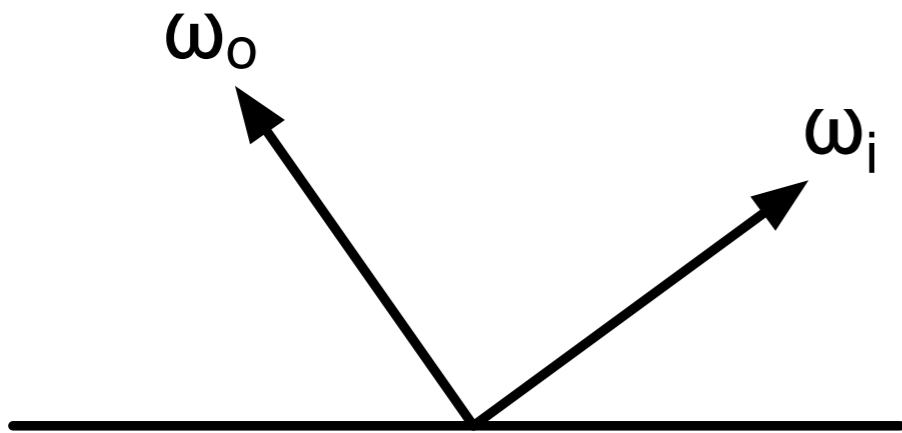
Overview

- **BRDFs**
- **Ambient Occlusion**
- **Spherical Harmonic lighting**
- Radiosity

BRDFs

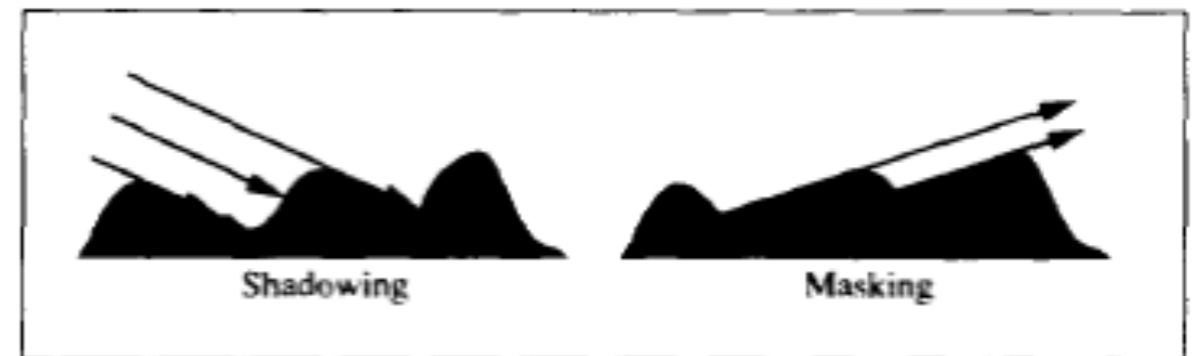
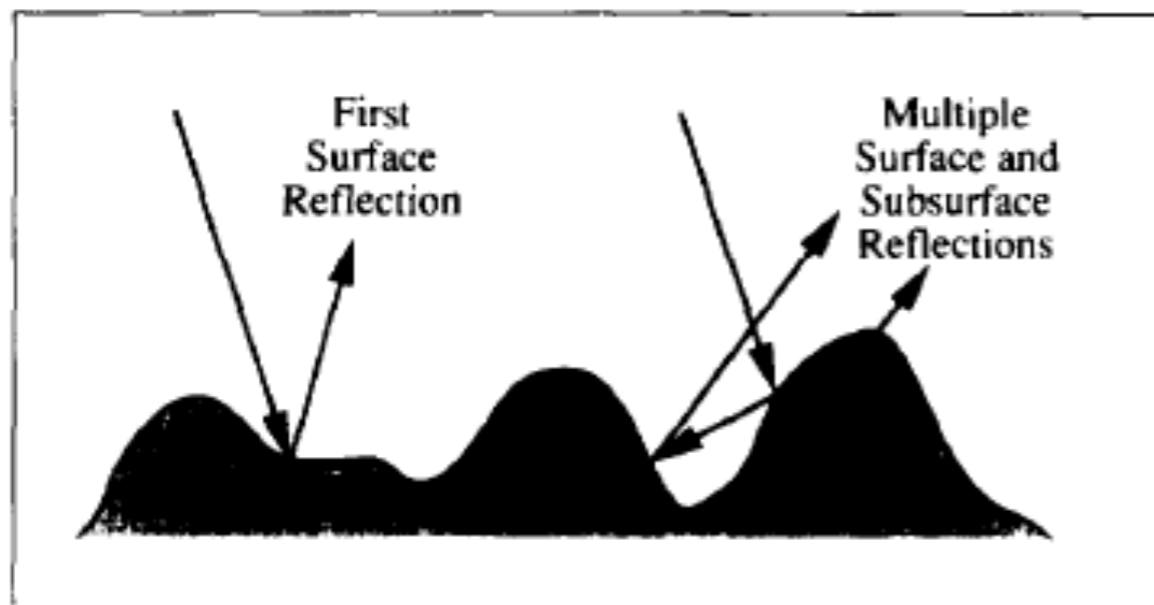
- Bidirectional reflectance distribution function
- A function of the incident and outgoing angles
- Can be dependent on position

$$f(x, \omega_o, \omega_i) = \frac{dL(\omega_o)}{dE(\omega_i)}$$



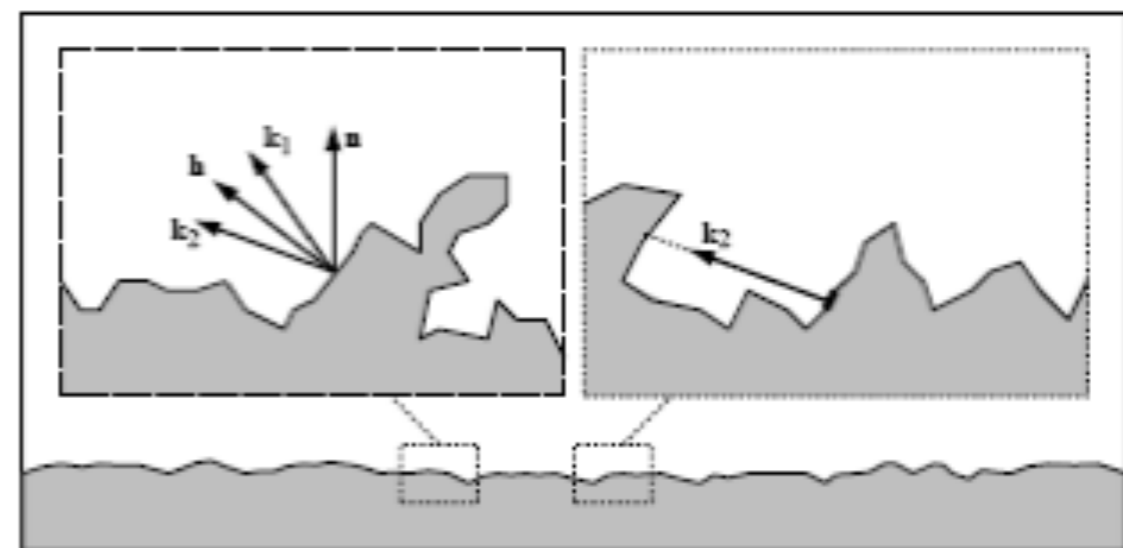
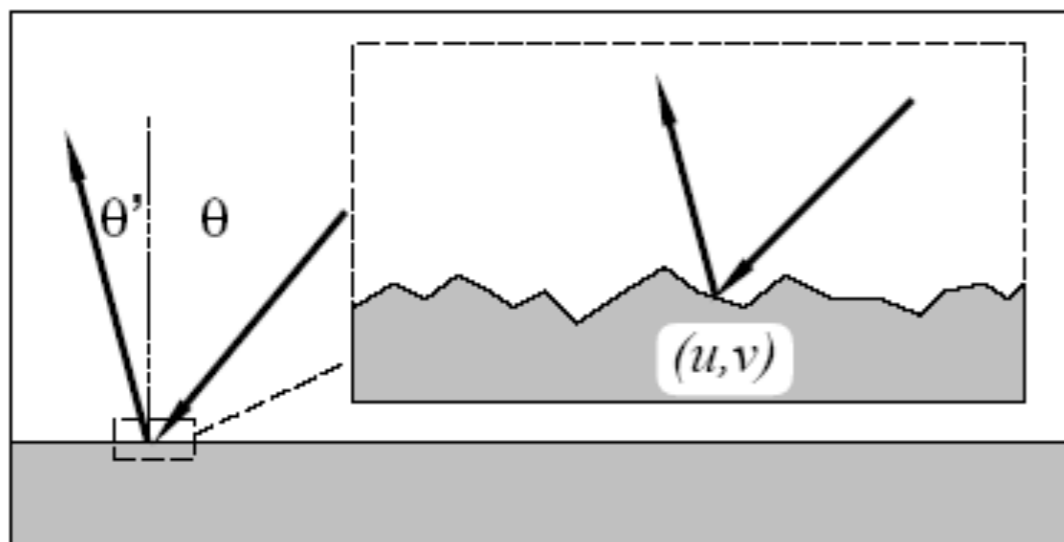
BRDFs

- Affected by the way light reflects on the surface
- Smoothness of surface:
 - Single reflection from a smooth surface - specular
 - Multiple reflections - diffuse
- Shadowing effect



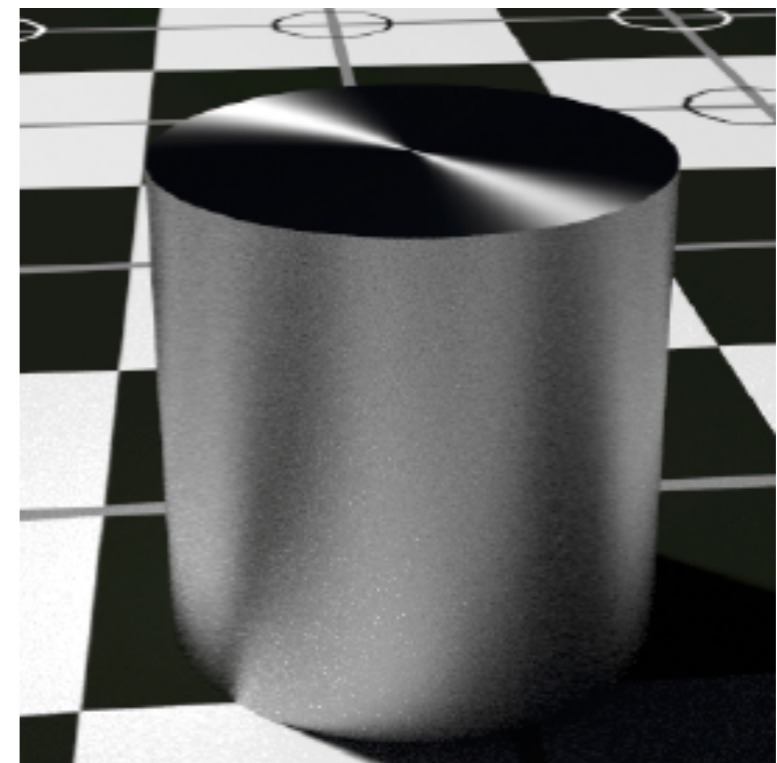
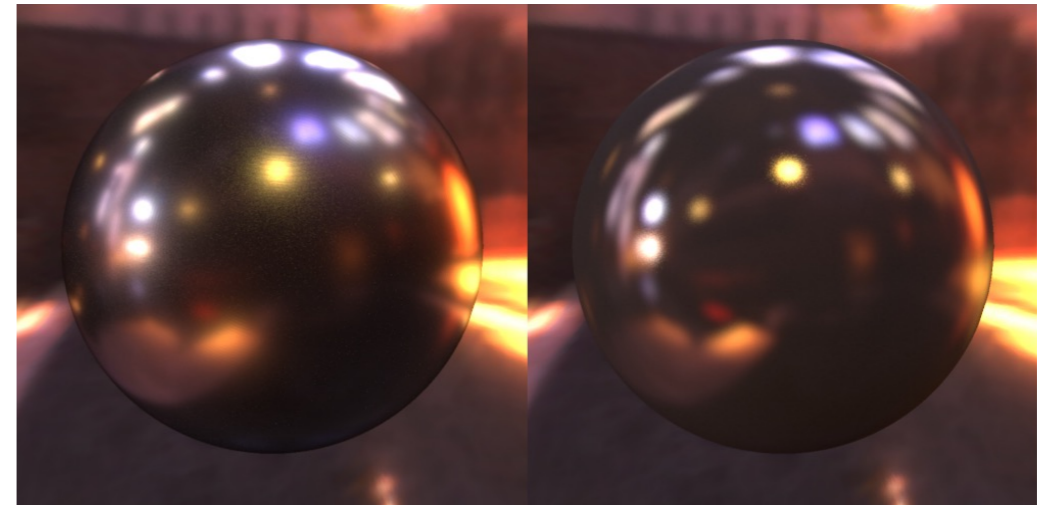
Microfacet theory

- [Torrance & Sparrow 1967]
 - Surface modeled by tiny mirrors
 - Value of BRDF at $\omega_i \omega_o$
 - # of mirrors oriented halfway between ω_o and ω_i
where ω_i is the incoming direction, ω_o is the out going direction
 - Modulated by Fresnel, shadowing/masking



Isotropic and anisotropic BRDFs

- Isotropic:
 - Lighting depends only on angle with normal and relative angle around normal
- Anisotropic:
 - Brushed metal
 - Can't be modelled using diffuse and specular reflection



Examples

- Ward

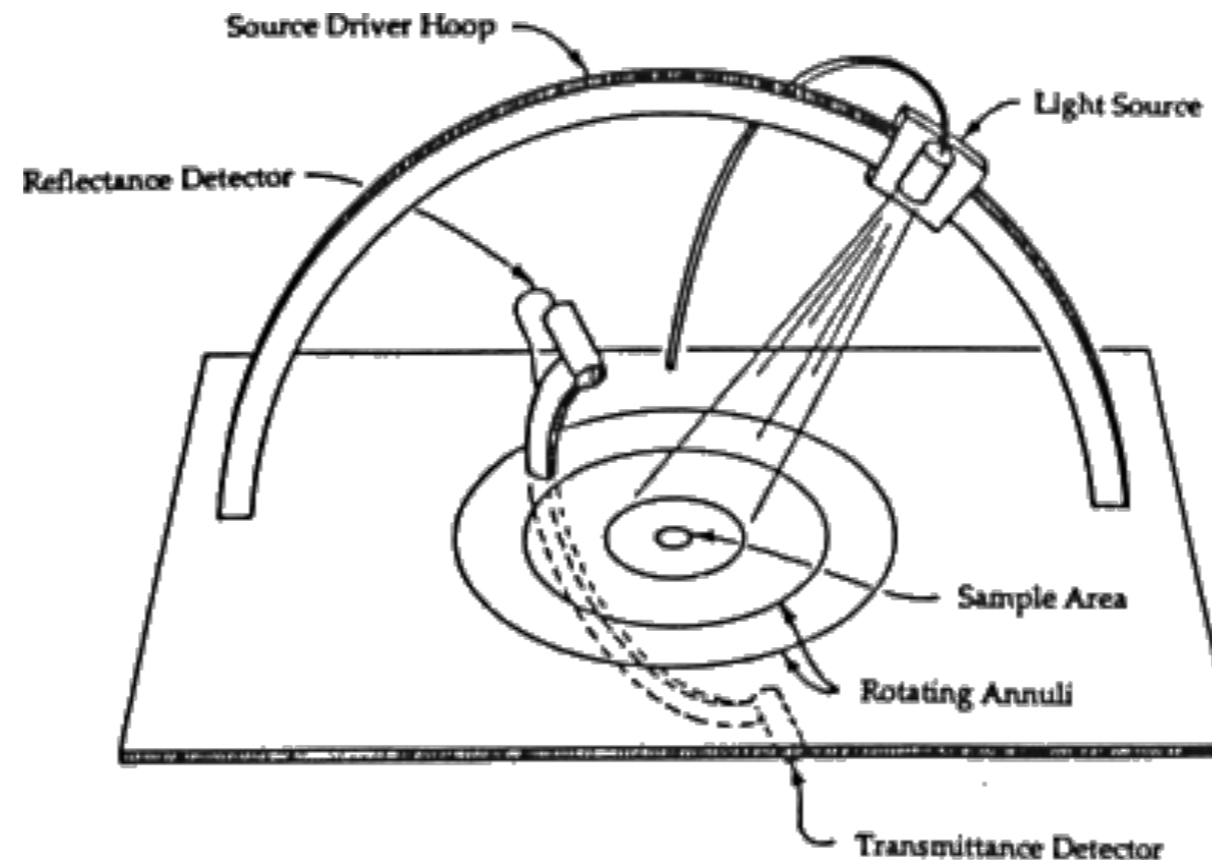
$$f(\omega_o, \omega_i) = \frac{\rho_s}{4\pi\alpha_x\alpha_y\sqrt{\cos\theta_i\cos\theta_o}} e^{\tan^2\theta_h\left(\frac{\cos^2\phi_h}{\alpha_x^2} + \frac{\sin^2\phi_h}{\alpha_y^2}\right)}$$

- Blinn-Phong:

$$f(\omega_o, \omega_i) = \frac{k_L}{\pi} + k_G \frac{8+s}{8\pi} (h \cdot n)^s$$
$$h = \frac{\omega_i + \omega_o}{2}$$

Capturing BRDFs

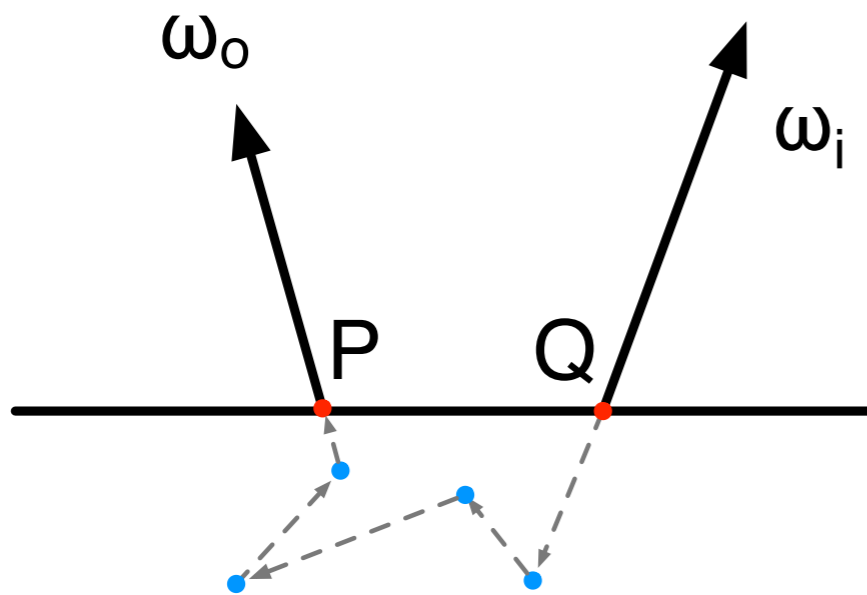
- Measured using a device called gonioreflectometer
 - Casting light from various directions onto the object, and capturing the light reflected back
- Problems:
 - Takes a long time and produces a huge amount of data (many GB)
 - Introduces errors



BSSRDF

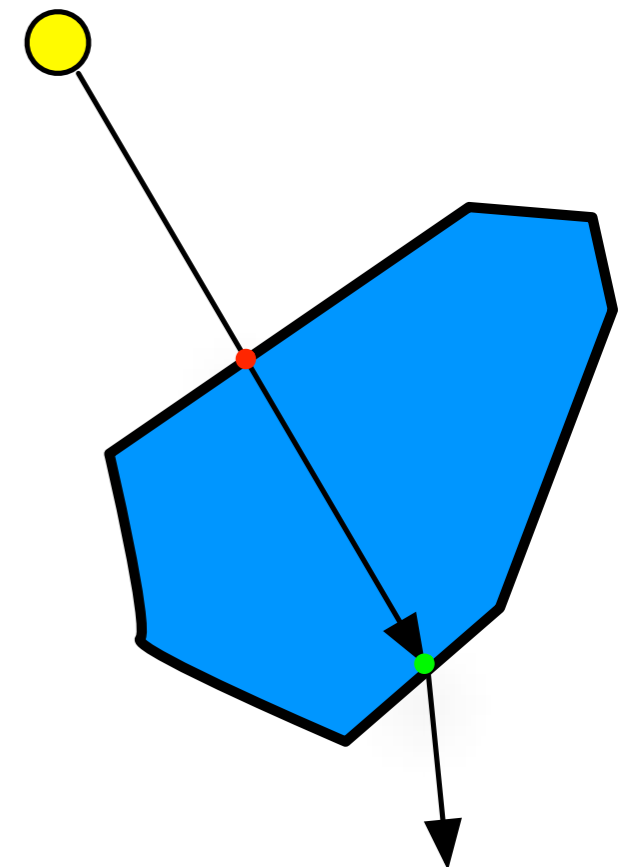
- Bidirectional surface scattering reflection distribution function
- Some surfaces exhibit subsurface scattering over a long range

$$f(P, Q, \omega_o, \omega_i) = \frac{dL(P, \omega_o)}{dE(Q, \omega_i)}$$



Rendering subsurface scattering

- For short distances, approximate with a BRDF
- Normal blurring
 - Use a bump map for specular reflection, but an unperturbed or blurred normal for diffuse shading
- Texture space diffusion
 - Render with diffuse lighting into a texture
 - Blur the texture and use for diffuse term in shading
- Depth map
 - Look up distance to surface in shadow map

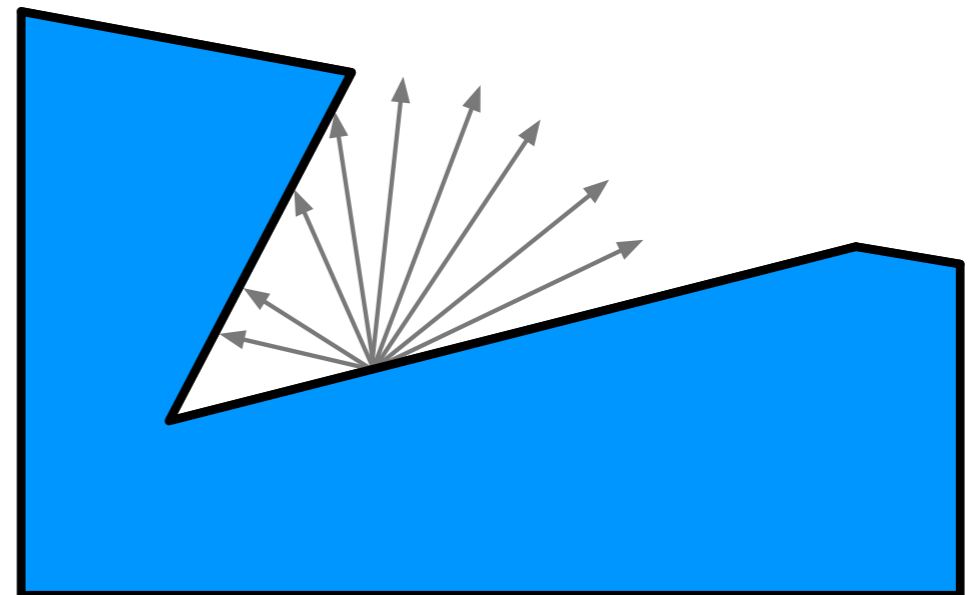
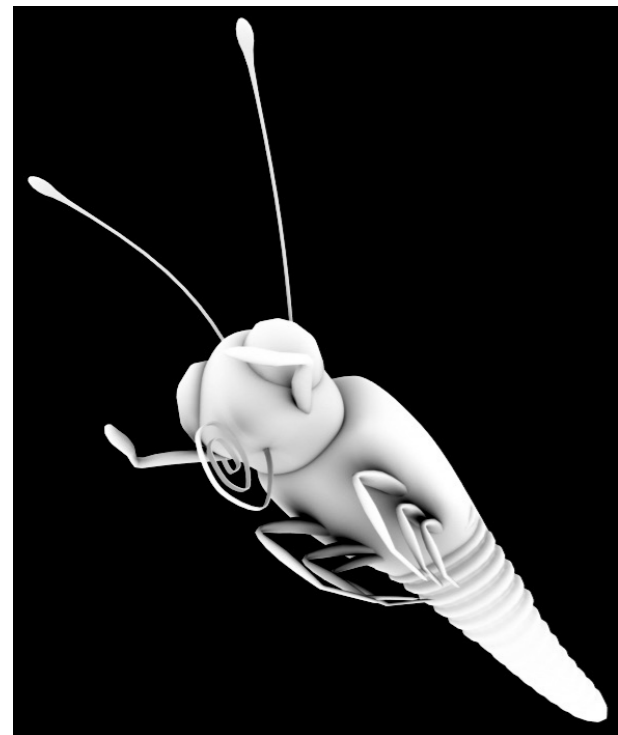


Ambient occlusion

- Shadowing of ambient light
- Incoming light is constant ambient lighting L from every direction, taking into account visibility function $v(x, \omega)$ at surface

$$E(x, n) = L_A \int_{\Omega} v(x, \omega_i) (n \cdot \omega_i) d\omega_i$$

$$k_A(x) = \frac{1}{\pi} \int_{\Omega} v(x, \omega) (n \cdot \omega_i) d\omega_i$$



Ambient occlusion

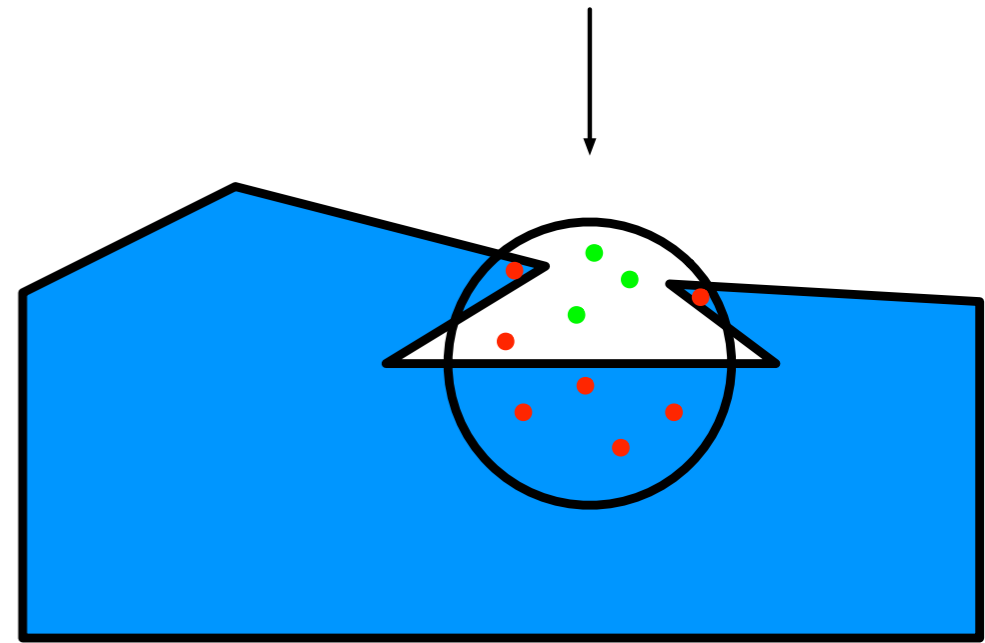
- This approach works well for discrete objects, what about closed rooms? Replace visibility function with a distance function.
- To combine with an environment map, we can apply normal bending:

$$n' = \frac{\int_{\Omega} v(\omega_i) \omega_i \cos \theta_i d\omega_i}{\left\| \int_{\Omega} v(\omega_i) \omega_i \cos \theta_i d\omega_i \right\|}$$

- Need to recalculate for non-static objects

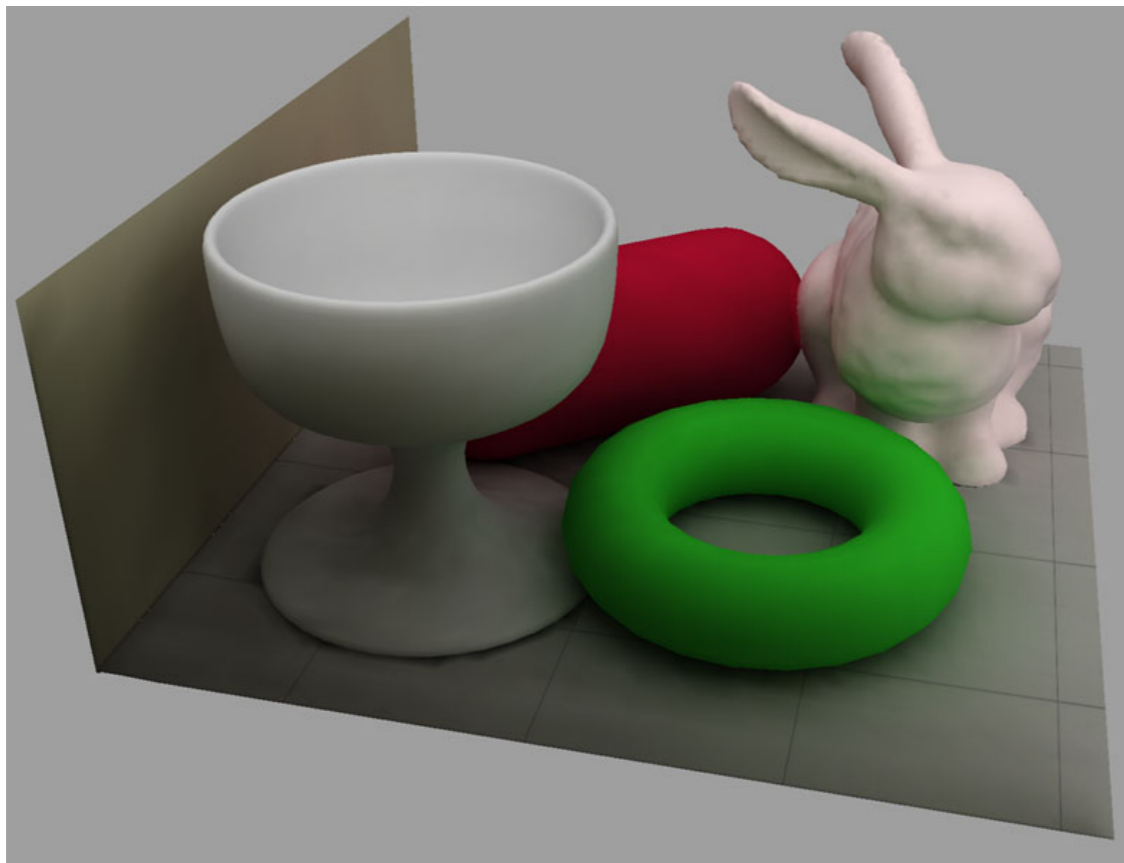
Screen space ambient occlusion

- Object space dynamic ambient occlusion is expensive
- Screen space ambient occlusion takes advantage of screen space depth buffer and surface normal information.
 - Sample points in a sphere around each pixel
 - Use fraction that are closer than Z-buffer depth value to estimate occlusion
 - Independent of object complexity

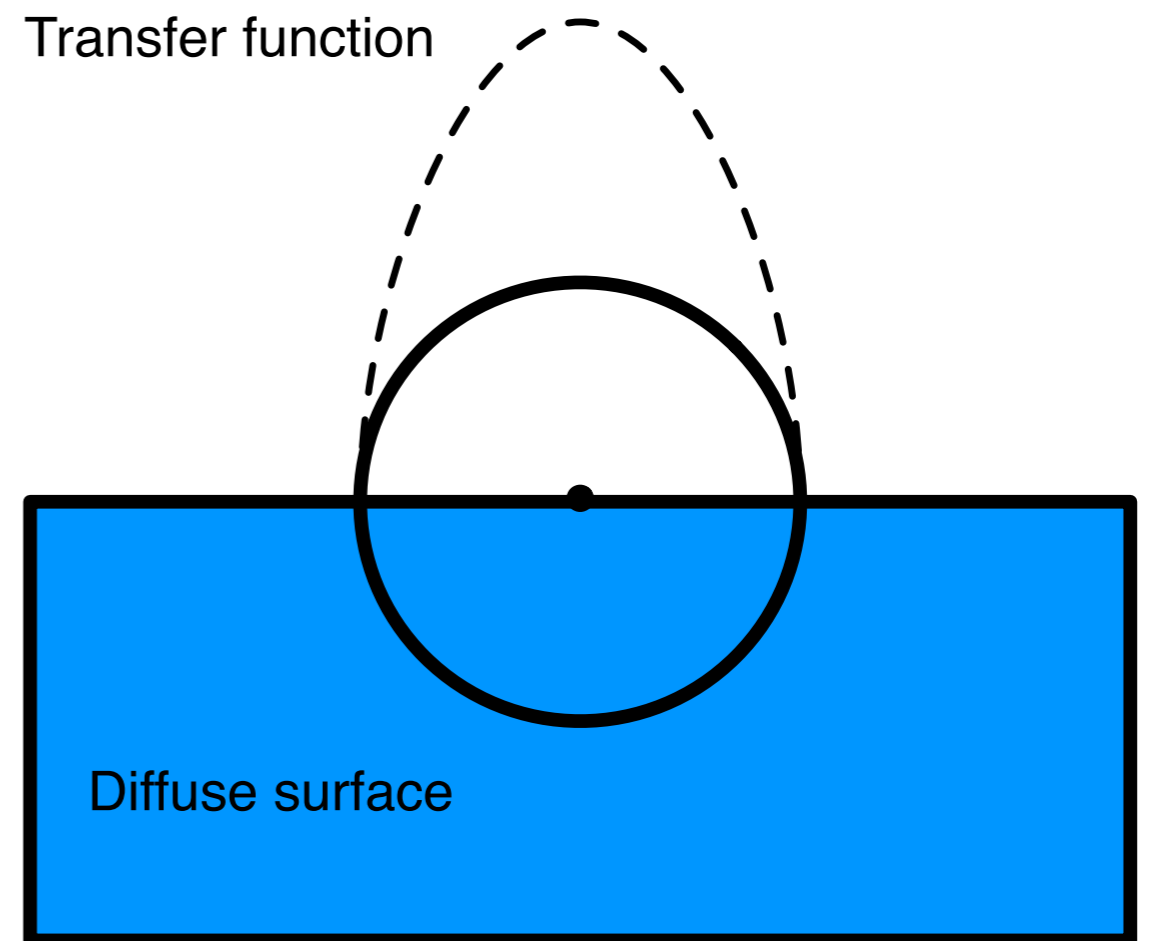
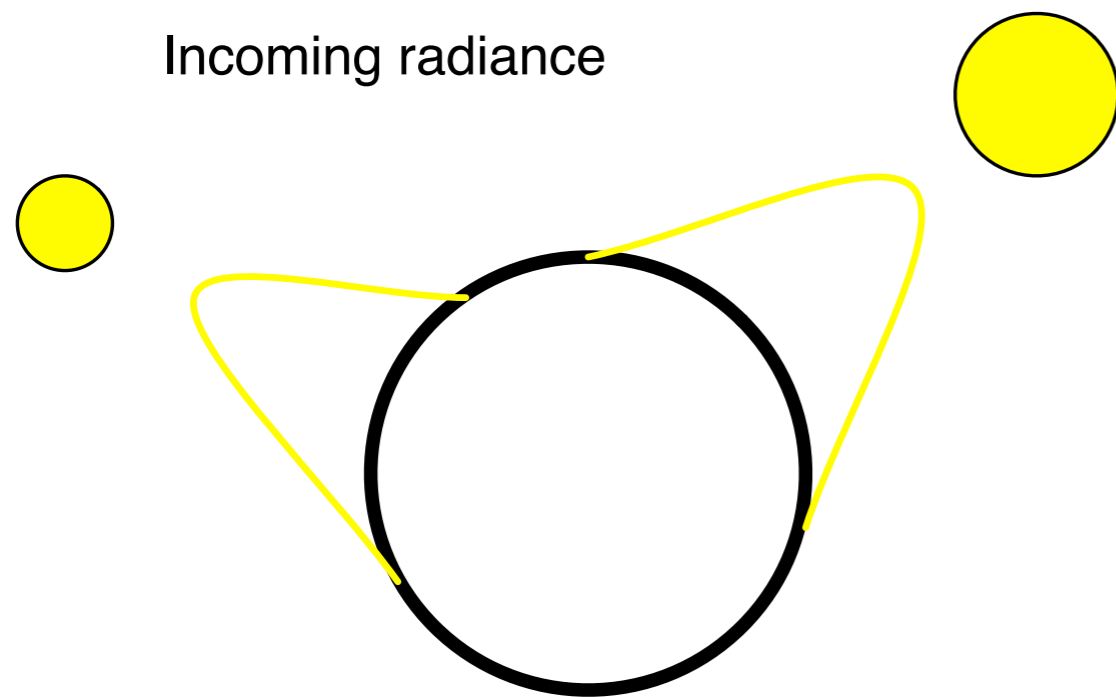


Spherical harmonic lighting

- Pre-computed radiance transfer (PRT) method
- Independent of view point, works with arbitrary lighting conditions

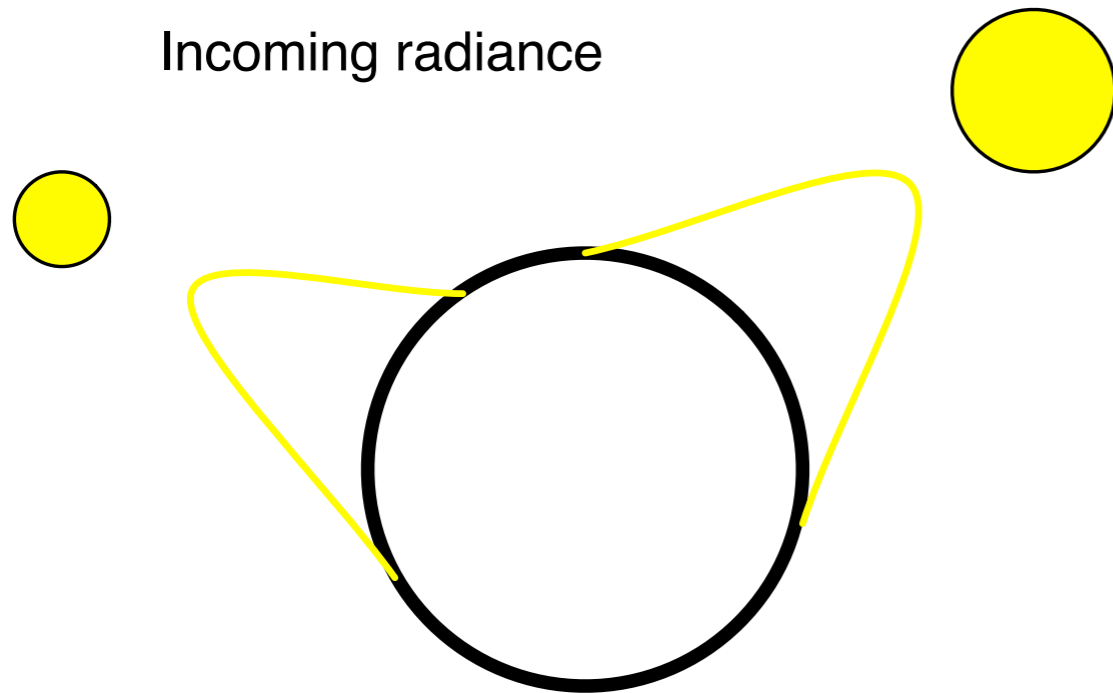


Overview

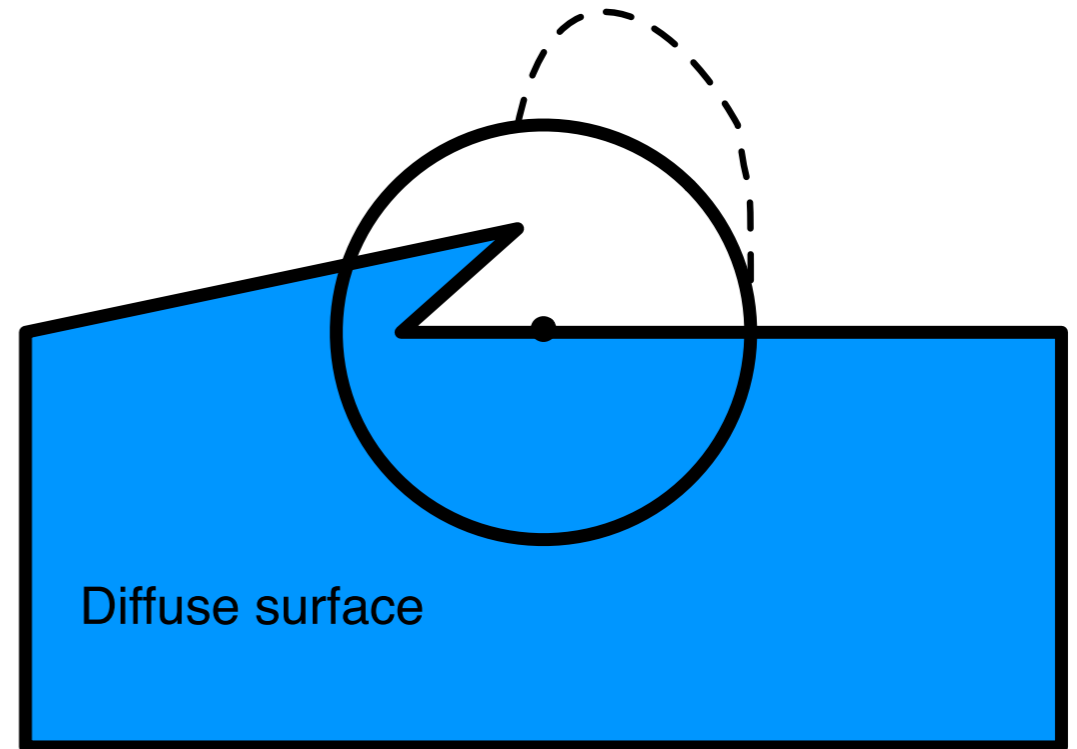


Overview

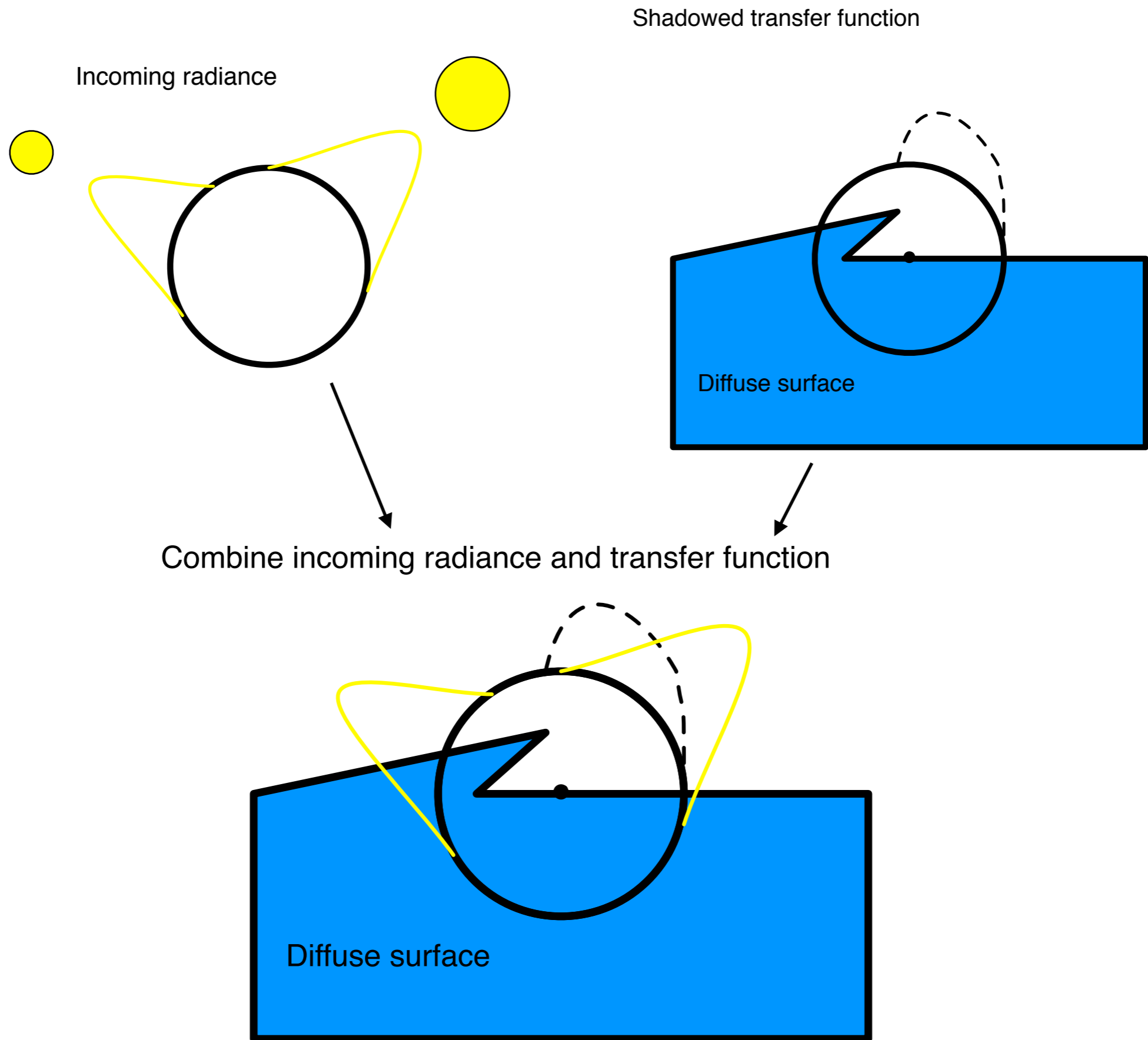
Incoming radiance



Shadowed transfer function

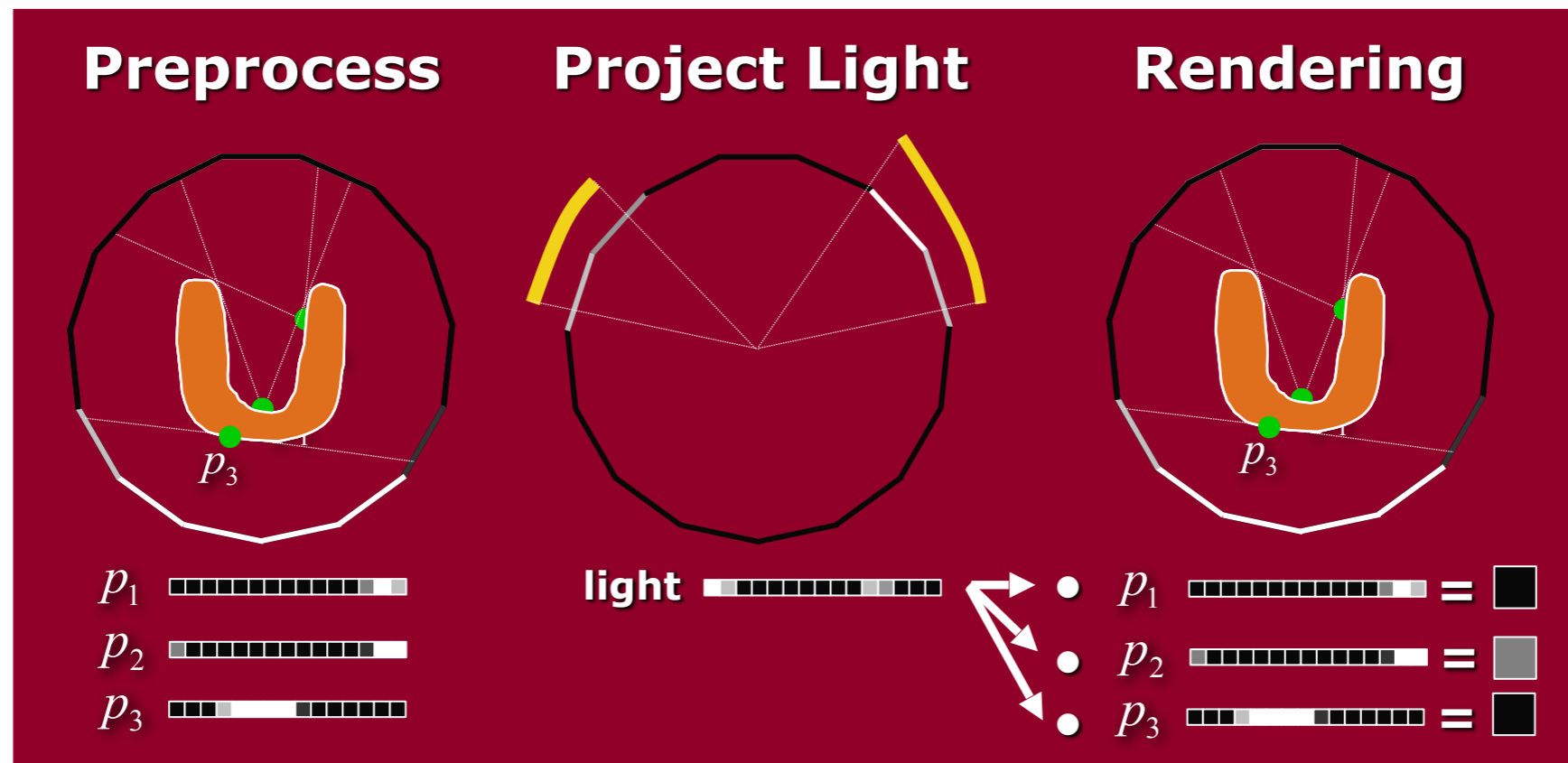


Overview



Diffuse light transfer

- 2D example, incoming directions form a circle. Only considering shadowing:



Spherical harmonics

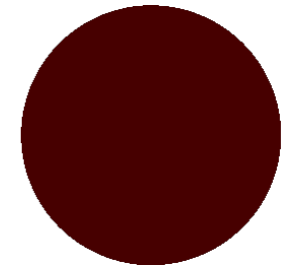
- Used as a basis to represent functions defined over a sphere

- Properties:

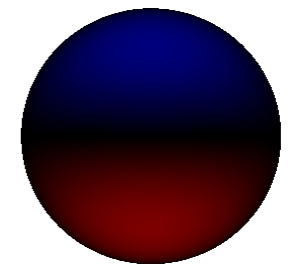
- Simple rotation and convolution

$$-l \leq m \leq l$$

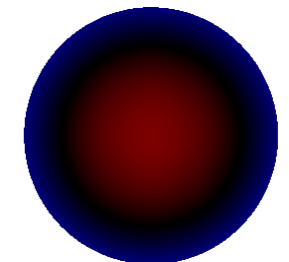
$$h_l^m(\theta, \phi) = \begin{cases} \sqrt{2}K_l^m \cos(m\phi) P_l^m(\cos \theta), & m > 0 \\ \sqrt{2}K_l^m \sin(-m\phi) P_l^{-m}(\cos \theta), & m < 0 \\ K_l^0 P_l^0(\cos \theta), & m = 0 \end{cases}$$



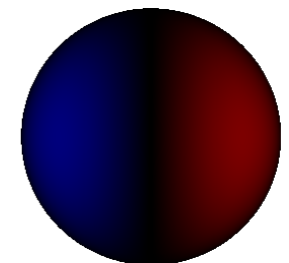
$l=0 \ m=0$



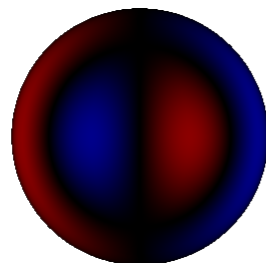
$l=1 \ m=-1$



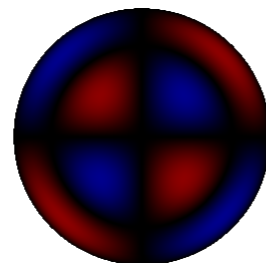
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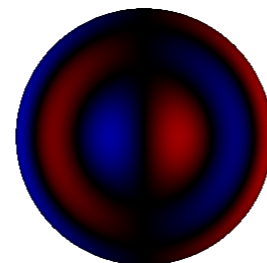
$l=1 \ m=1$



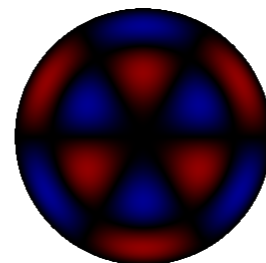
$l=2 \ m=1$



$l=3 \ m=-1$



$l=3 \ m=2$



$l=4 \ m=-2$

Spherical harmonic lighting

- Diffuse lighting (constant for every outgoing direction) based on incoming light and transfer function

$$\int_{\Omega} L_i(\omega) t(\omega) d\omega$$

- Calculate integral by multiplying SH coefficients

$$\int_{\Omega} L_i(\omega) t(\omega) d\omega = \sum_j^{\infty} c_{L,j} c_{t,j}$$

- Approximation based on a finite number of coefficients

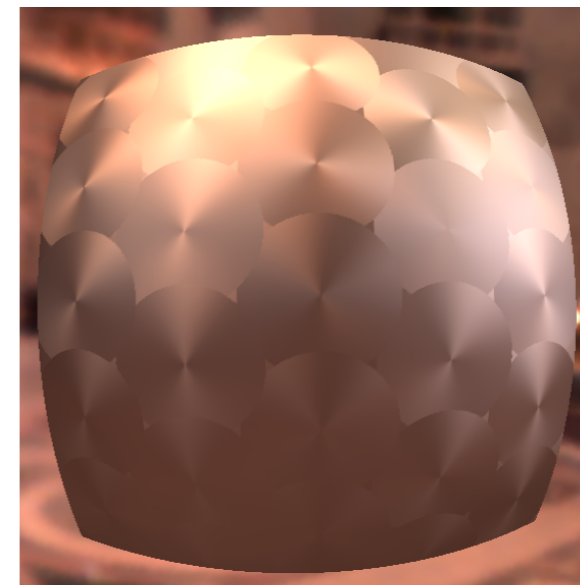
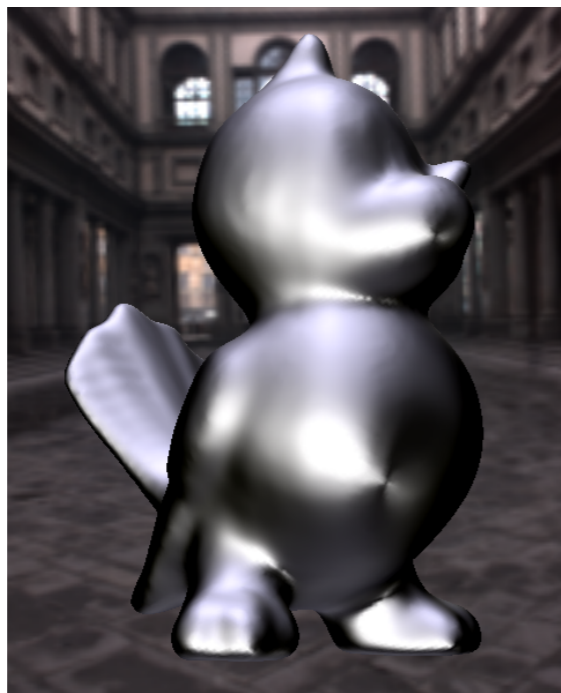
Examples



Shadows

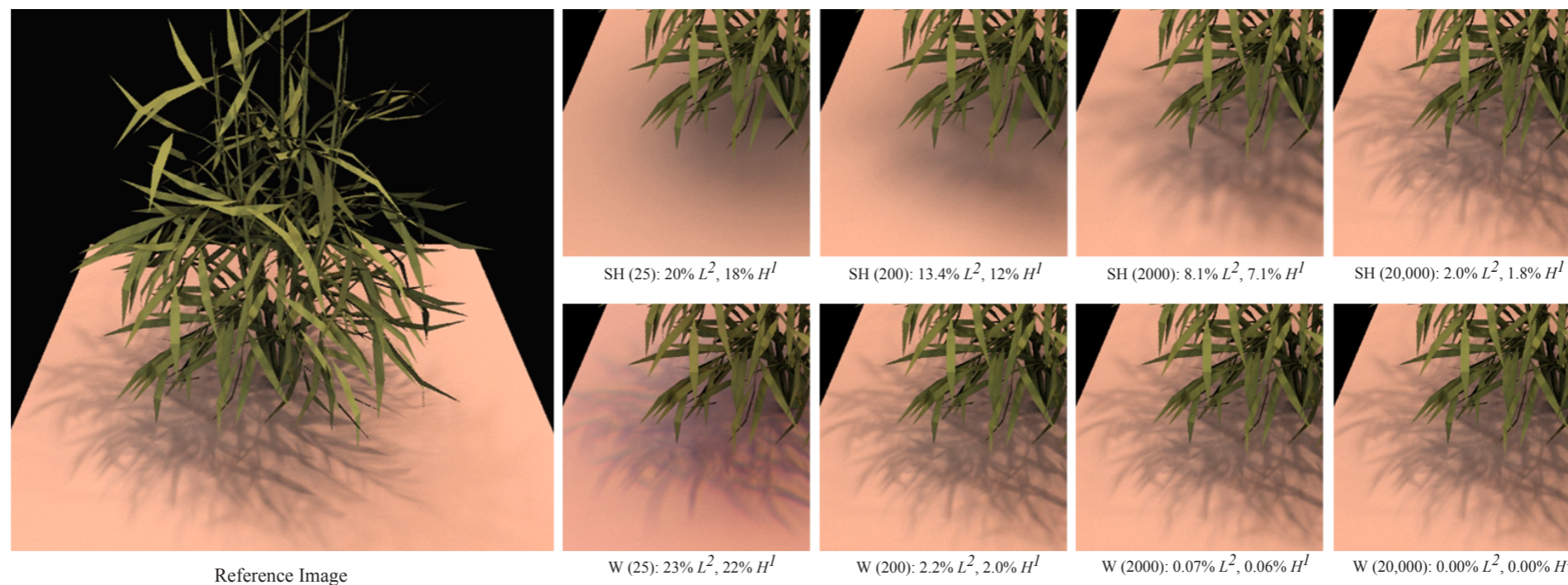
1 bounce
interreflections

2 bounce
interreflections



Problems

- Only valid for static objects
- To represent sharp changes (e.g. shadows) or high frequency lights we need many spherical harmonic coefficients
- An alternative approach - Haar wavelets



Overview

- BRDFs
- Ambient Occlusion
- Spherical Harmonic lighting
- **Radiosity**

Radiosity rendering

- Based on a method developed by researchers in heat transfer in 1950s
- Applied to computer graphics in the mid 1980s
- by Michael Cohen Tomoyuki Nishita



(a)



Radiosity rendering

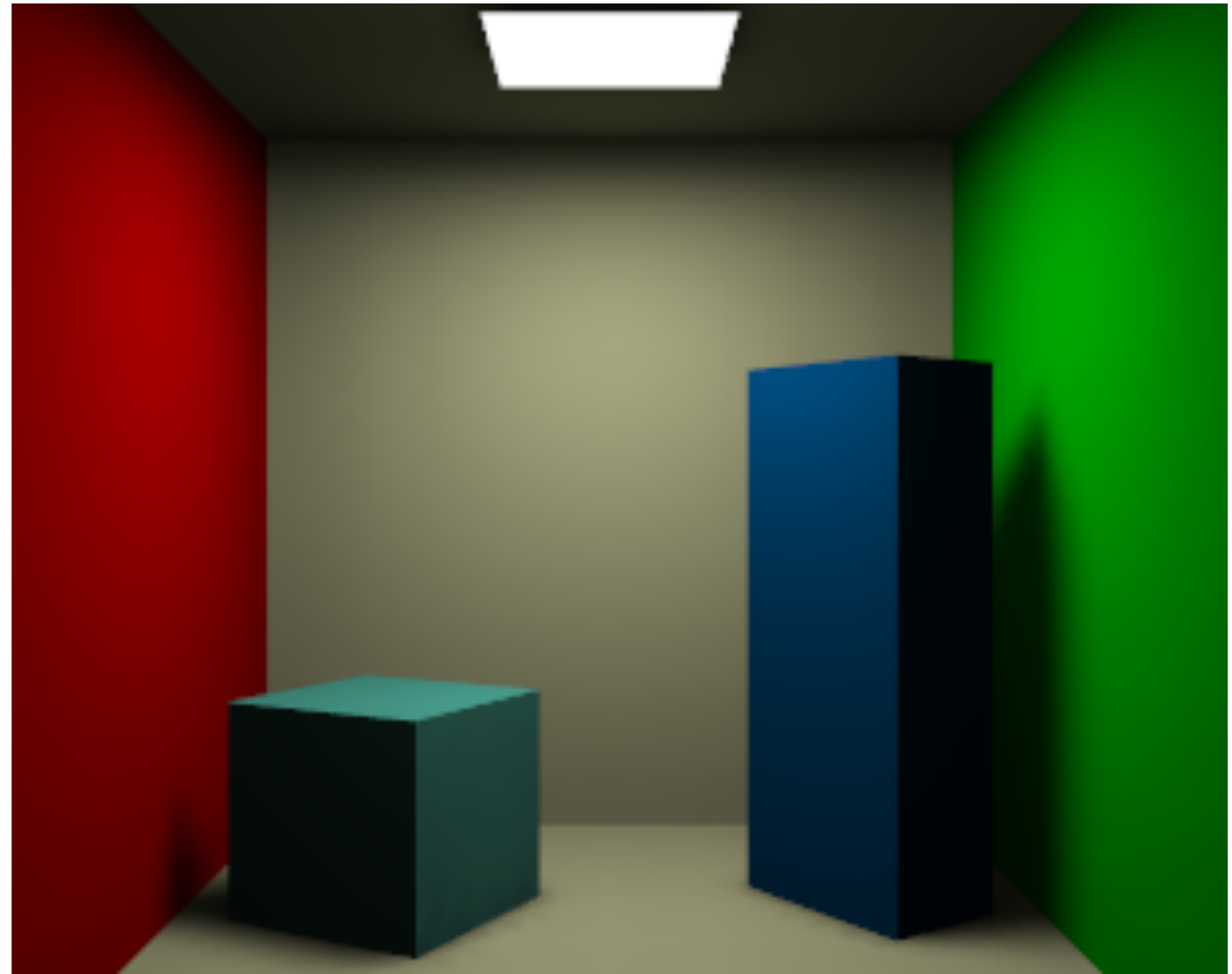
- Can simulate inter-surface reflection
- Can produce nice ambient effects
- Can simulate effects such as
 - Soft shadows,
 - color bleeding
- Can only handle diffuse colour
- → need to be combined with ray-tracing to handle specular light



(a)

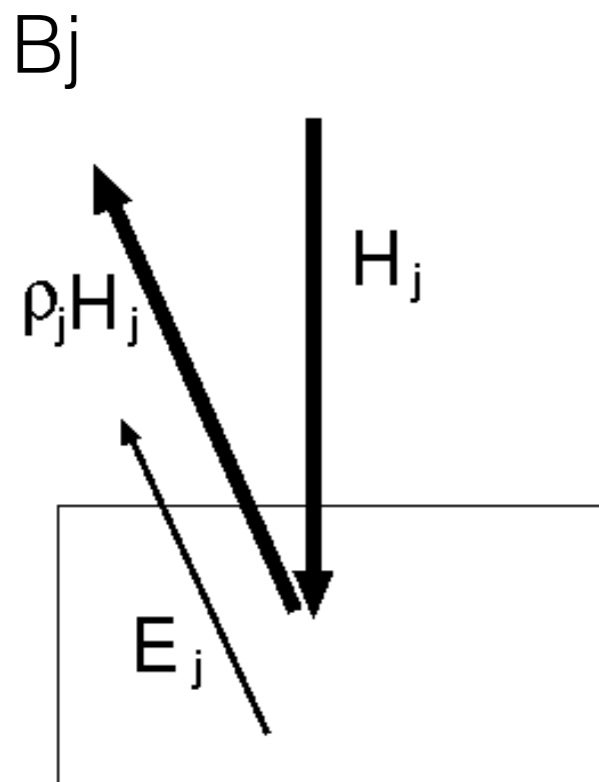


Colour bleeding and soft shadows



The Radiosity model

- At each surface in a model the amount of energy that is given off (Radiosity) is comprised of
 - the energy that the surface emits internally (E), plus
 - the amount of energy that is reflected off the surface (ρH)



$$B_j = \rho_j H_j + E_j$$

B_j is the radiosity of surface j ,

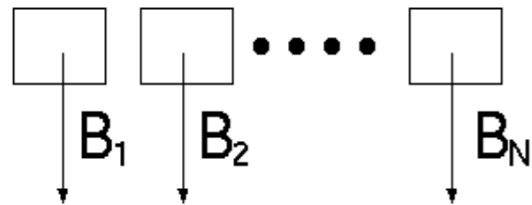
ρ_j is the reflectivity of surface j ,

E_j is the energy emitted by surface j .

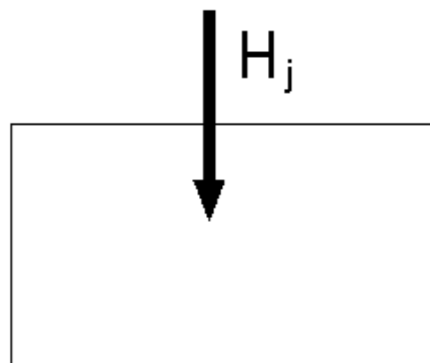
H_j is the energy incident on surface j

The Radiosity model

- The amount of incident light hitting the surface can be found by summing for all other surfaces the amount of energy that they contribute to this surface



$$H_j = \sum_{i=1}^N B_i F_{i,j}$$



$F_{i,j}$ Form factor

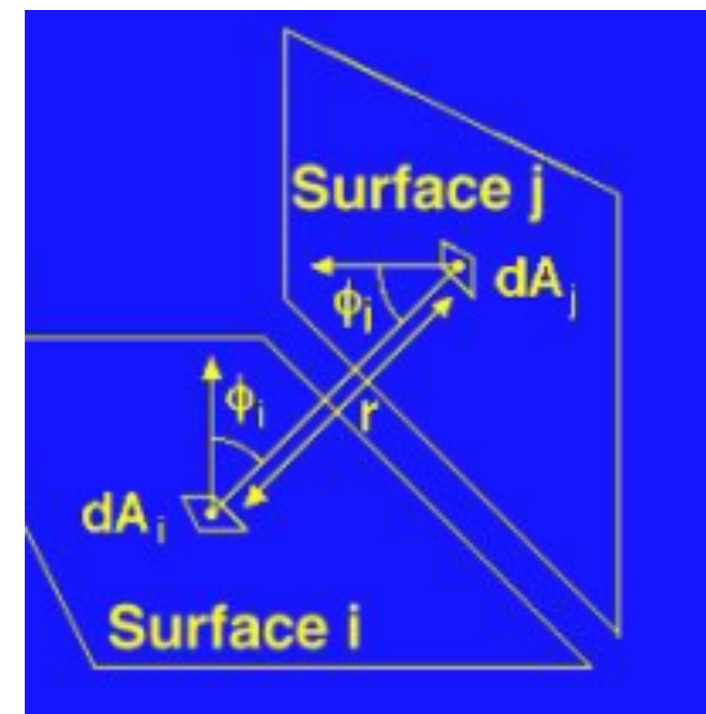
Form factors

- The fraction of energy that leaves surface i and lands on surface j
- Between differential areas, it is

$$\frac{\cos \phi_i \cos \phi_j}{\pi |r|^2} dA_i dA_j$$

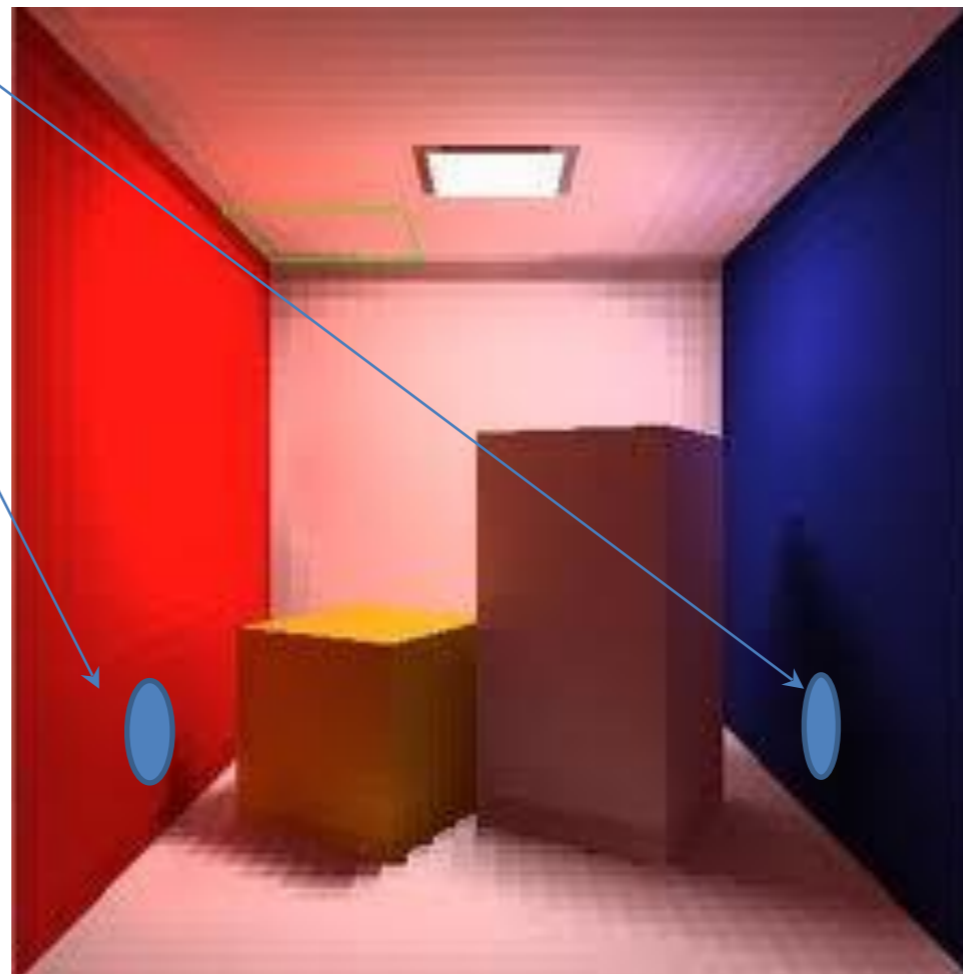
- The overall form factor between i and j is

$$F_{ij} = \sum_i \sum_j \frac{\cos \phi_i \cos \phi_j}{\pi |r|^2} dA_i dA_j$$



Form factors

- Also need to take into account occlusions
- The form factor for those faces which are hidden from each other must be zero



Radiosity matrix

- The radiosity equation now looks like this:

$$B_j = E_j + \rho_j \sum_{i=1}^N B_i F_{i,j}$$

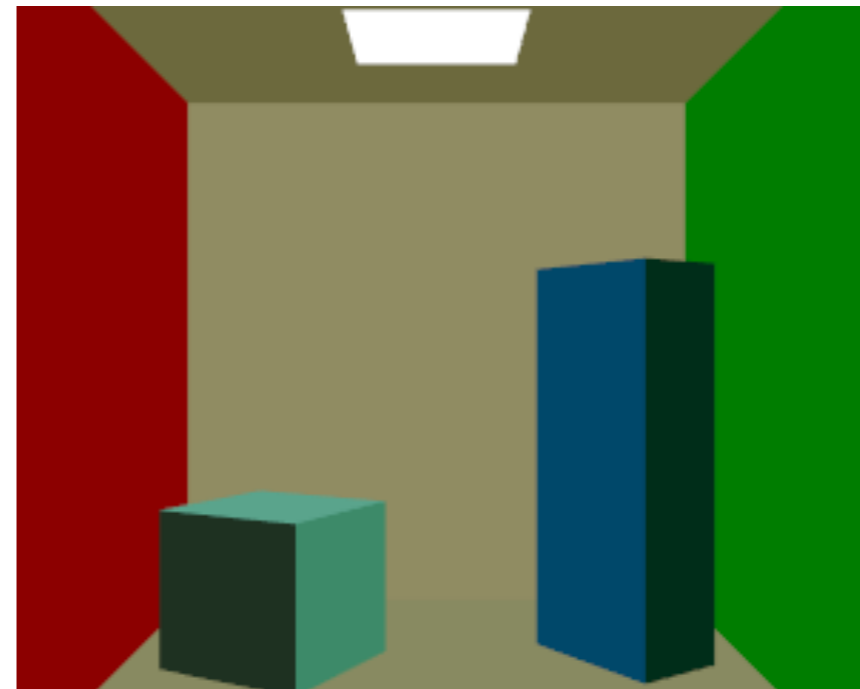
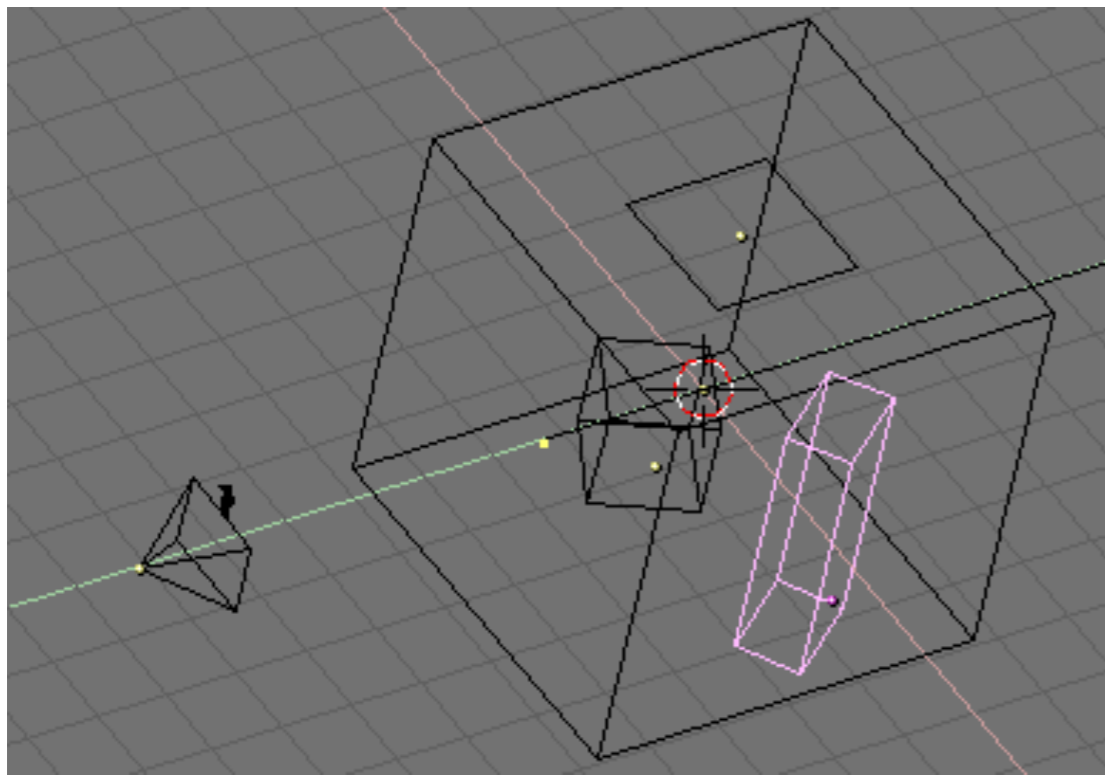
- The derived radiosity equations form a set of N linear equations in N unknowns. This leads nicely to a matrix solution:

$$\begin{pmatrix} 1 - \rho_1 F_{11} & -\rho_1 F_{12} & \dots & -\rho_1 F_{1N} \\ -\rho_2 F_{21} & 1 - \rho_2 F_{22} & \dots & -\rho_2 F_{2N} \\ \vdots & \vdots & \dots & \vdots \\ -\rho_N F_{N1} & -\rho_N F_{N2} & \dots & 1 - \rho_N F_{NN} \end{pmatrix} \begin{pmatrix} B_1 \\ B_2 \\ \vdots \\ B_N \end{pmatrix} = \begin{pmatrix} E_1 \\ E_2 \\ \vdots \\ E_N \end{pmatrix}$$

- Solve for B – Use methods like Gauss-Seidal

Radiosity steps

1. Generate Model (set up the scene)
2. Compute Form Factors and set the Radiosity Matrix
3. Solve the linear system
4. Render the scene



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Radiosity steps

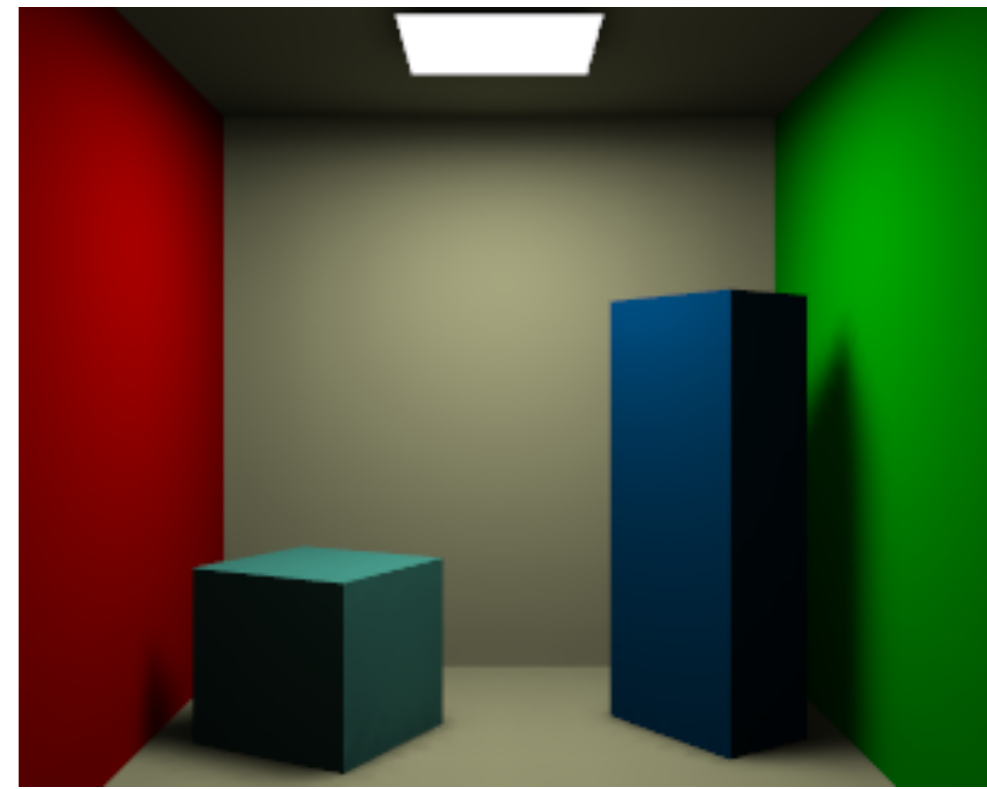
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Radiosity steps

1. Generate Model
2. Compute Form Factors and set the Radiosity Matrix
3. Solve the linear system
4. Render the scene

$$\begin{pmatrix} 1 - \rho_1 F_{11} & -\rho_1 F_{12} & \dots & -\rho_1 F_{1N} \\ -\rho_2 F_{21} & 1 - \rho_2 F_{22} & \dots & -\rho_2 F_{2N} \\ \vdots & \vdots & \dots & \vdots \\ -\rho_N F_{N1} & -\rho_N F_{N2} & \dots & 1 - \rho_N F_{NN} \end{pmatrix} \begin{pmatrix} B_1 \\ B_2 \\ \vdots \\ B_N \end{pmatrix} = \begin{pmatrix} E_1 \\ E_2 \\ \vdots \\ E_N \end{pmatrix}$$

- Where do we resume from if objects are moved?
- Where do we resume from if the lighting is changed?
- Where do we resume from if the reflectance parameters of the scene are modified?
- Where do we resume from if the view point changes?

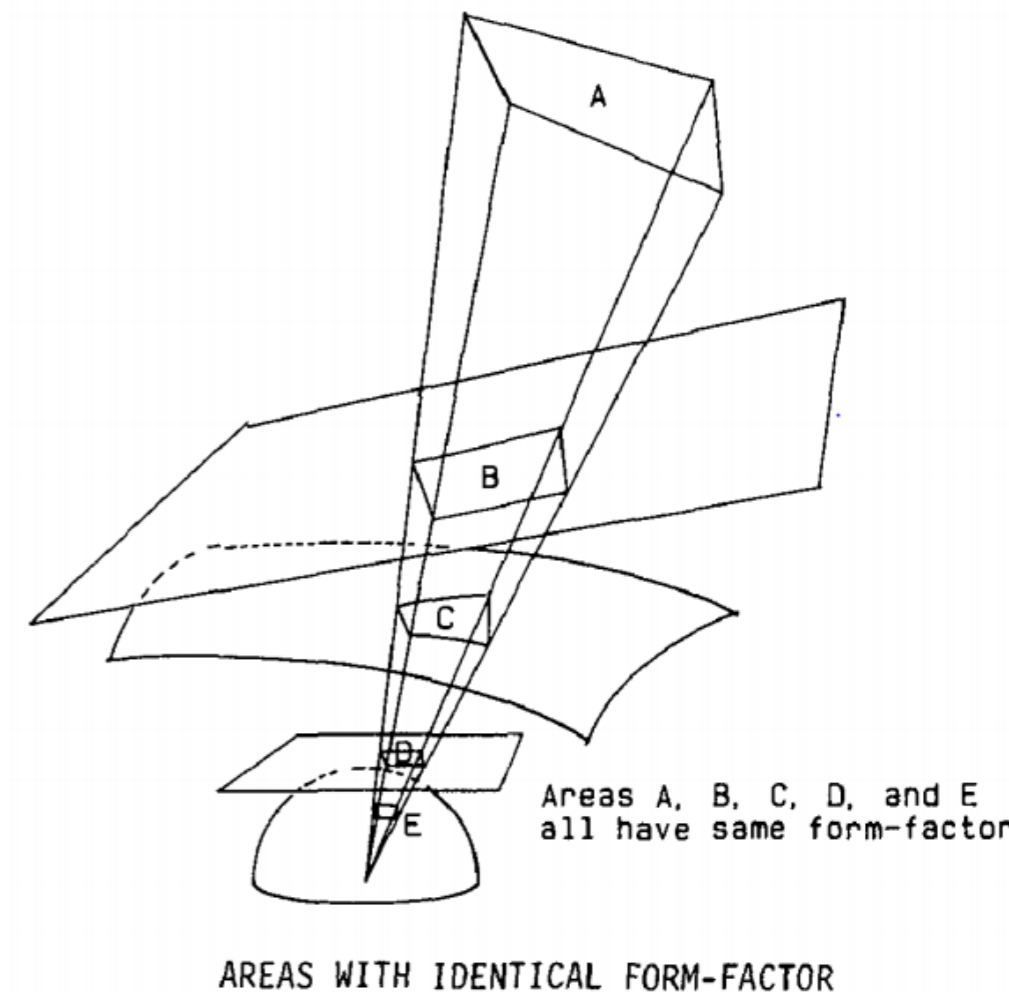
Radiosity features

- Very costly
- The faces must be subdivided into small patches to reduce the artefacts
- The computational cost for calculating the form factors is expensive
 - Quadratic in the number of patches
- Solving for B_i is also very costly
 - Cubic in the number of patches
- Cannot handle specular light

Hemicube method

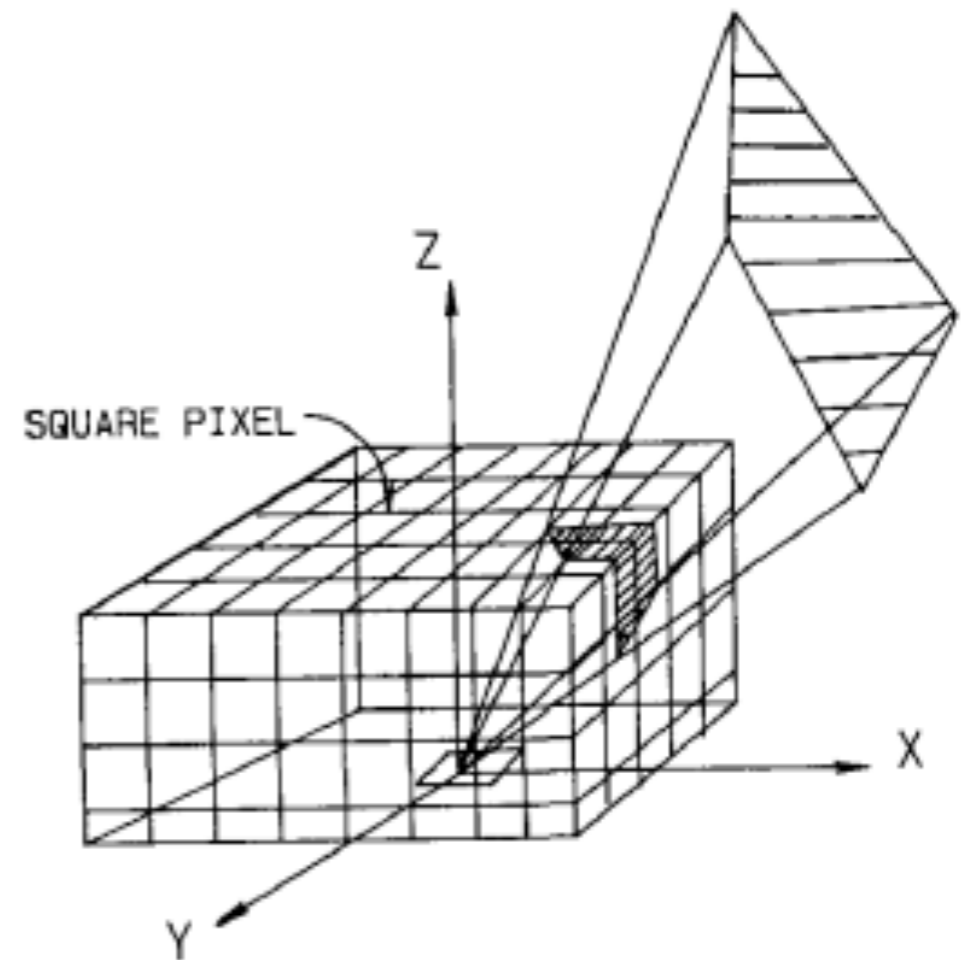
by Michael Cohen'85

- Accelerates the computation of the form factor
- The form factor for the right four faces with respect to a small patch in the bottom is the same
- Then, we can project the patches onto a hemicube



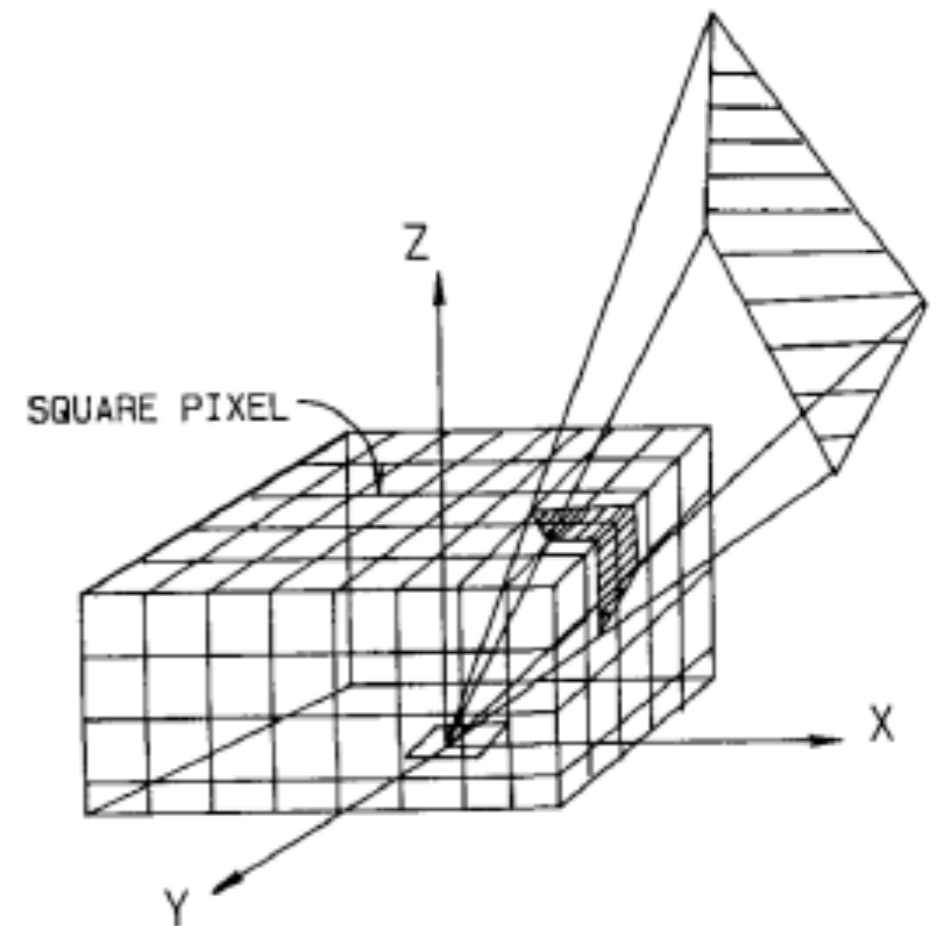
Hemicube method

- Prepare a hemicube around the patch i
- Project those polygons you want to compute the form factor with patch i onto the hemicube
- Then, compute the form factor between them



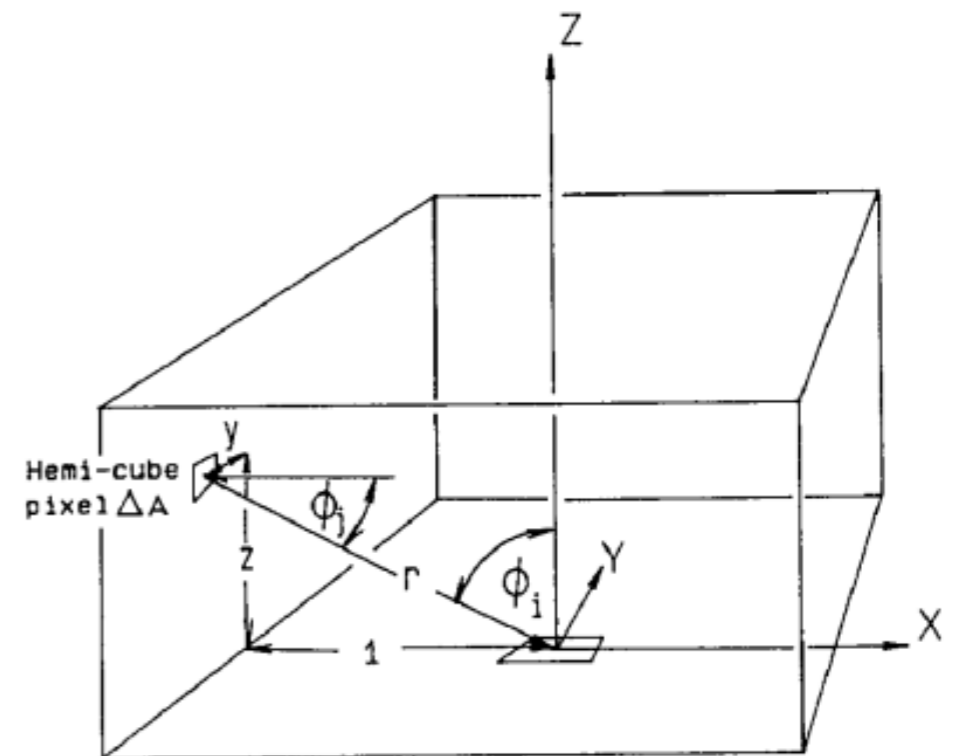
Hemicube method

- This can be done by perspective projection
- We can use the Z-buffer algorithm to find the closest polygon
- Handling the occlusion
- Setting the form factor of the pairs that occlude each other to zero



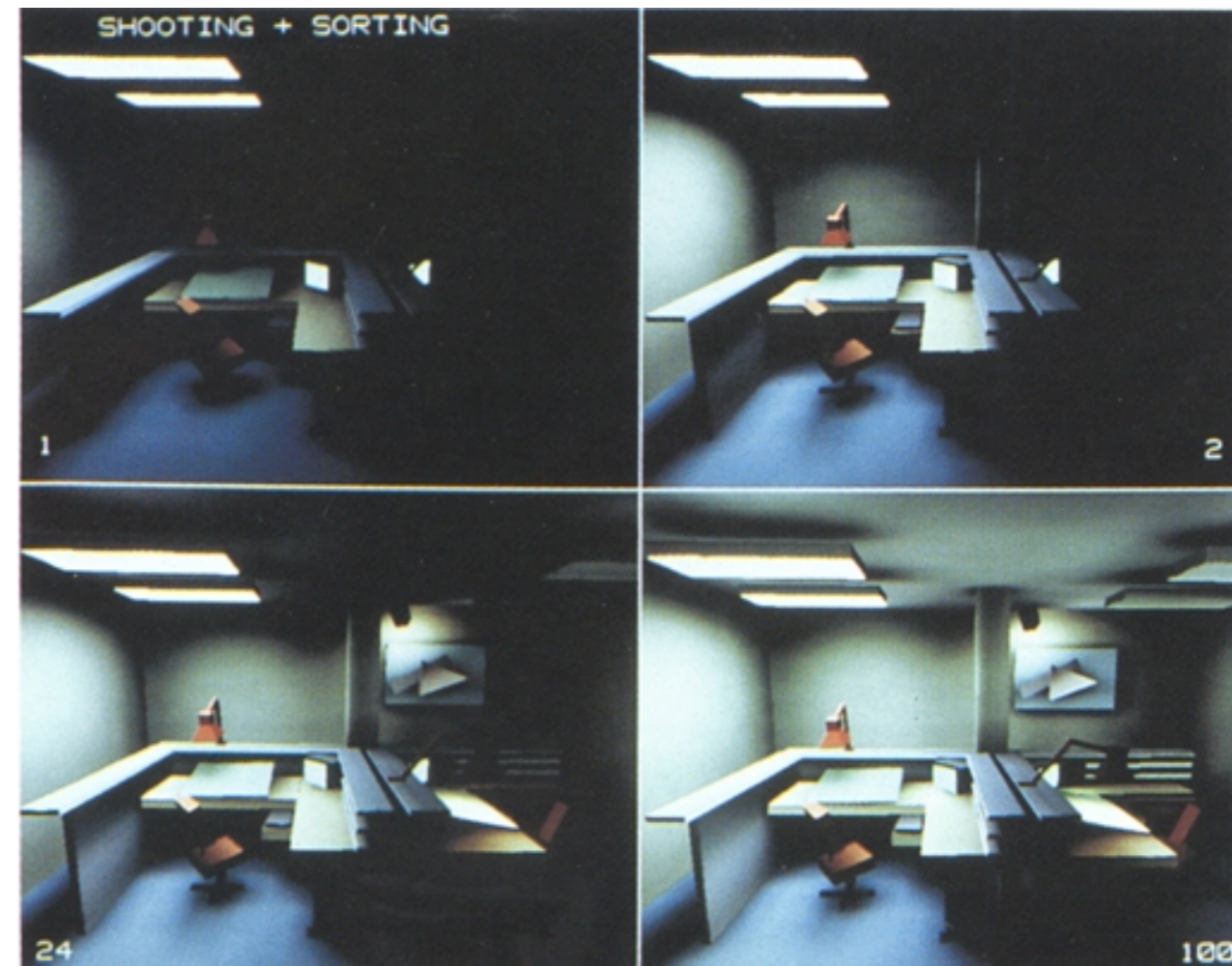
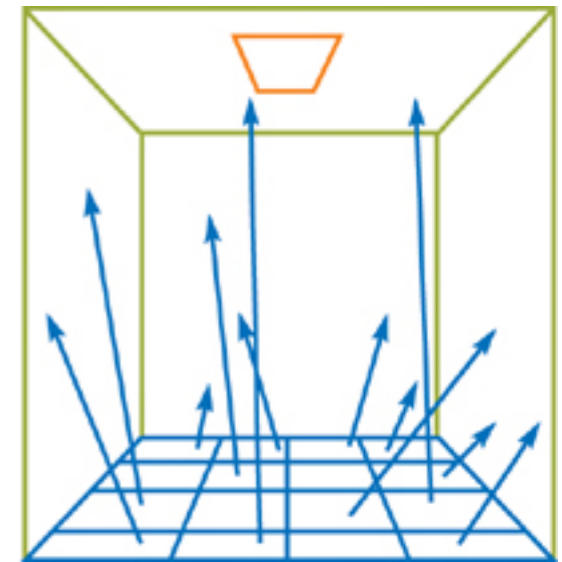
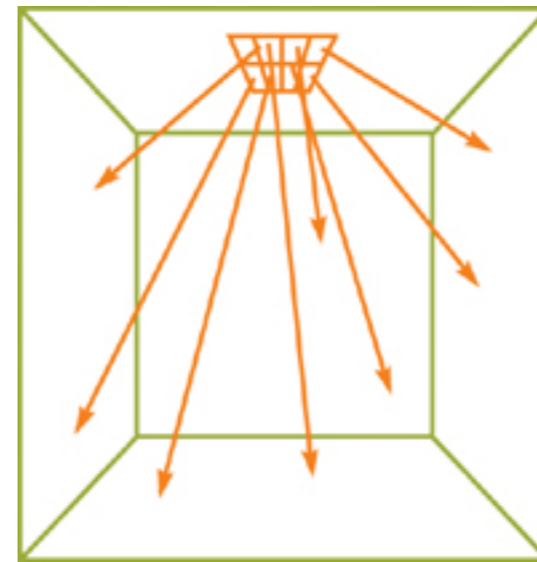
Hemicube method

- The form factor between each pixel of the hemicube and the patch at the origin can be pre-computed and saved in a table
- Only 1/8 of all are needed, thanks to symmetry



Progressive refinement

- Find patch in scene with largest undistributed radiosity and distribute it into the scene
- Store radiosity shot into scene and mark as undistributed
- Iterate until converged
- No need to store radiosity matrix (form factors computed on the fly)



Summary

- BRDFs describe how a surface reflects light
- Ambient occlusion approximates shadowing of ambient lighting
- Spherical harmonic lighting allows us to efficiently render pre-calculated radiance transfer under arbitrary lighting conditions
- Radiosity can simulate diffuse inter-reflectance
- Form factor computation can be accelerated by hemi-cube.

References

- BRDFs: Shirley Chapter 20.1.6
- Ambient occlusion: Akenine-Möller Chapter 9.2
- Spherical harmonic lighting: Akenine-Möller Chapter 9.11
- Radiosity:
 - Foley Chapter 16.13
 - Extra: http://http.developer.nvidia.com/GPUGems2/gpugems2_chapter39.html

References - papers

- Spherical harmonics:
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